Heat Problems Associated with Genesis Flood Models—Part 1: Introduction and Thermal Boundary Conditions

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Abstract

The Genesis Flood must have produced drastic geological changes involving extremely energetic processes which must also have generated an enormous heat load. Yet since the inhabitants of the Ark and many aquatic creatures survived the Flood, the heat must have been removed without raising temperatures beyond biological endurance limits. This is the first in a series of papers intended to identify and, where possible, to quantify the most important sources of Flood heat. Here we review the thermal state of the earth's mantle, crust and oceans, both now and in the recent geological past, and take note of mantle heterogeneities as tracers of its tectonic history. Extensive temperature records based on the oxygen-18 content of marine fossil shells indicate geologically recent ocean floor temperatures not exceeding 13°C; all models of the Flood and its aftermath need to take this limit into account. A recent attempt to model ocean floor cooling within a biblical chronology by invoking a transient subsurface heat sink along with rapid oceanic plate movement was unsuccessful, suggesting that important hitherto-neglected geophysical effects may need to be included in future. Inventories of heat-producing radioactive elements estimated from rock compositions and from geoneutrino measurements are in broad agreement, lending confidence to the anticipated process of estimating the heat load due to accelerated nuclear decay during the Flood. Preliminary estimates for granite suggest that this radiogenic heat load would be overwhelming without the operation of a powerful insitu cooling mechanism which has yet to be identified.

Keywords: Mantle heterogeneities; temperatures; oxygen isotope palaeothermometry; oxygen-18; heat flow; radiogenic heat

Introduction

Chapters 7 and 8 of Genesis describe the catastrophic global Flood which God brought about at the time of Noah as judgment upon an incurably wicked human race and our environment. This Flood was of such scale and intensity that it must have drastically reshaped the face of the earth and deposited enormous quantities of heat. Since Noah, his family, the animals on the Ark, and many aquatic creatures survived the Flood, the atmosphere and oceans could not have been heated beyond biological endurance limits. The resulting scientific problem is to explain how the heat was removed without raising environmental temperatures too high.

This series of articles seeks to identify and, where possible, to quantify the sources of Flood-related heat in order to provide boundary conditions and guidelines for creation scientists seeking suitable explanations of how Flood and post-Flood environmental temperatures were kept within limits. Since some of the key background literature arises from the RATE (Radioisotopes and the Age of The Earth) project, the relevant RATE results are reassessed, partly to place on record an independent appraisal of RATE and some of the ensuing published exchanges.

This article (Part 1) deals with boundary conditions relevant to modeling the earth's thermal history. These include its internal temperature field, mantle structure, past and present ocean temperatures, surface heat flows and its inventory of heat-producing radionuclides. Given that several ocean temperature indicators are in current use, we focus here on ocean temperatures reconstructed from oxygen-18 data. Part 2 deals with newer, less familiar indicators including minor and trace element ratios, biomolecular indices and clumped isotopes. Part 3 reviews vapour canopy models, which are typically characterized by high pre-Flood atmospheric temperatures and a large heat load produced by collapse of the canopy at the beginning of the Flood. Part 4 assesses the heat released in Catastrophic Plate Tectonics (CPT), currently one of the most successful approaches to Flood modeling, and in similar Flood models; here the heat problem is caused by the large quantity of hot mantle material surfacing in a short time. Part 5 addresses probably the most acute Flood-related heat source, Accelerated Nuclear Decay (hereafter AND), which has been invoked to explain several key RATE findings; this heat source was recognized by the RATE authors (Vardiman et al. 2005). Part 5 also reassesses the methods and results of three components of RATE, viz. radiohalos in granites, fission tracks in zircons and helium diffusion in zircons.

Several authors have suggested that during the Flood the earth must have suffered a major bombardment from space; the resulting surface heat load is assessed in Part 6. Part 7 summarizes

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our conclusions and suggests possible directions for future investigations which might reveal a solution to Flood heat problems or to clarify key related questions.

Interior Temperatures and Structure of the Earth

Present-day conditions within the earth have been modeled using seismic, gravity, and geomagnetic data combined with compositional modeling of its constituent materials. Compositional models are informed by comparison with other bodies in the Solar System (sun, meteorites, and other planets). A representative model is the PREM (Preliminary Reference Earth Model) of Dziewonski and Anderson (1981); the data of interest here is conveniently summarized by Stacey and Davis (2008; Appendix F1 lists mechanical properties, Appendix G thermal properties). The resulting global averages as functions of distance r (in km) from the center of the earth are shown in fig. 1. The CMB (core-mantle boundary) shows its presence by a jump in temperature and a bend in the pressure profile, indicating a jump in density, at a depth of 2891 km (r=3480 km). Thermal boundary layers are evident (1) at the CMB, and (2) in the upper mantle between the transition zone (whose upper boundary is at a depth of 410 km; see Frost 2008; Stacey and Davis 2008) and the surface. Such boundary layers are expected in the presence of the mantle convection associated with plate tectonic motions because there can be no advective heat transport through the CMB or through the lithosphere except at MORs and other localized volcanic outlets. The temperature drop through the region above the transition zone is approximately 1600 K. The larger part of this occurs across the lithosphere, where also the temperature gradient is steepest; the thickness of the lithosphere, although very variable (Paysanos 2008), is typically of order 100km.



Fig. 1. Earth's globally averaged internal pressures (black) and temperatures (red) as functions of radius (r), based on data from Stacey and Davis (2008).

Given spherically symmetric models of the kind proposed by Dziewonski and Anderson (1981) as a reference, geophysical evidence, mainly seismic, reveals considerable mantle heterogeneity (Kennett and Tkalčić 2008). The heterogeneity is especially pronounced in the D" region, the lowest 200–300 km immediately above the CMB, which is characterized seismically by a decreasing gradient of both P-wave and S-wave velocities with depth (Loper and Lay 1995). The heterogeneity is understood to be in temperature, composition, and material phase (Čížková et al. 2010; Hirose and Lay 2008; Klein, Jagoutz, and Behn 2017).

Regions of increased P- and S-wave speeds generally signify descending slabs, interpreted as subducted material, while lower wave speeds signify upwellings or mantle plumes. In many cases subducted material has apparently been flattened in the transition region around 660km below the surface (Fukao, Widiyantoro, and Obayashi 2001). However in the Farallon slab under North and Central America and the Indian/Tethys slab under the Himalayas and Bay of Bengal, subduction has reached the lower mantle. Baumgardner (2003) cites as evidence of mantle heterogeneity: (1) a ring of dense material at the bottom of the mantle, lying roughly below the perimeter of today's Pacific Ocean; (2) superswells (e.g. in the South Pacific and under the East Africa Rift Valley), regions of locally elevated surface caused by rising mantle plumes (Grand, van der Hilst, and Widiyantoro 1997). Baumgardner (2003) cites density homogeneities up to the 3–4% level, and assuming this to be entirely thermal in origin, he infers temperature differences of up to 3000-4000K in the underlying mantle. He argues that differences this large are unlikely to have persisted for 100Ma (million years) as in the uniformitarian chronology, but are perfectly plausible if the subduction occurred only a few thousand years ago. However Baumgardner does not attempt to justify his assumption that compositional and phaserelated contributions to density can be ignored.

Mantle heterogeneities, especially where they relate to subducted material, thus serve as an archive of earth's tectonic history (Kennett and Tkalčić 2008). For example, a recent study has revealed, in uniformitarian terms, a link between the flux of subducted material reaching the CMB and the timing of geomagnetic field reversals (Hounslow, Domeier, and Biggin 2018). Mantle heterogeneities and the structures they represent should therefore be taken into account in the development of coherent, fully integrated Flood models; Baumgardner (2003) has considered them qualitatively in this light. However a detailed assessment of the significance of mantle heterogeneities in the context of Flood heat and of Flood modeling more generally is beyond the intended scope of this article; such an exercise should be undertaken in due course.

Ocean Temperatures

Considerable geological evidence is available relating to the earth's past ocean temperatures and ice cover. The temperature data comes from marine deposits from the whole span of the geological record, notably from Phanerozoic fossil shells. The record is particularly rich for ocean-floor sediment thanks to extensive ocean floor drilling programs carried out over the last 50 years, viz. the International Ocean Discovery Program: Exploring the Earth Under the Sea (IODP 2017) and its earlier incarnations. The resulting data is publicly accessible via the PANGAEA® online database (PANGAEA 2017). Geologically most ocean floors are identified as Jurassic, Cretaceous, or Cenozoic, i.e. not exceeding 200 Ma in uniformitarian terms, although the Ionian Sea and the Eastern Mediterranean are dated from about 270-230Ma (Müller et al. 2008). In the CPT model of Austin et al. (1994), in which the Flood/ post-Flood boundary is tentatively located at the K-Pg (Cretaceous-Palaeogene) interface, these stratigraphic assignments correspond to mid-Flood to post-Flood times. In other Flood models the correspondence of ocean floor development with Flood chronology varies according to where the Flood/post-Flood boundary is placed; see Section 3 for further discussion of alternative Flood models.

The most widely used temperature and ice-cover indicator is the oxygen isotope ratio (¹⁸O/¹⁶O) in marine fossil shells and in ice: hence the term *oxygen isotope paleothermometry*. This has been complemented over the last 20–30 years by other temperature indicators including fossil shell Mg/Ca ratios, coral Sr/Ca ratios, biomolecular indicators and "clumped isotopes," which are all reviewed in the next paper (Part 2) in this series.

1. Principles of Oxygen Isotope Paleothermometry

Oxygen has three stable isotopes, ¹⁶O, ¹⁷O, and ¹⁸O, with respective molar abundances on earth of 99.757(16)%, 0.038(1)%, and 0.205(14)% (De Laeter et al. 2003); the bracketed numbers are the quoted uncertainties in the last figures. The ¹⁸O/¹⁶O ratio in any given sample is expressed via the δ^{18} O value, defined by

$$\delta^{18} O = \left(\frac{R_{sample}}{R_{standard}} - 1\right) \times 1000 \tag{1}$$

where R represents the ¹⁸O/¹⁶O ratio and δ ¹⁸O values are quoted in parts per thousand (per mil), signified by the symbol ∞ . The standard for δ ¹⁸O values for marine fossil shells is known as VPDB

(Vienna PeeDee Belemnite) and for water (ice, snow, liquid, or vapour) VSMOW (Vienna Standard Mean Ocean Water); there is a small offset between these standards (USGS 2004; Ravelo and Hillaire-Marcel 2007). The long-term reproducibility of the laboratory standard for carbonate δ^{18} O values is within ±0.08‰ (Ravelo and Hillaire-Marcel 2007).

Oxygen isotope paleothermometry is founded on "the temperature dependence of oxygen isotope fractionation between authigenic minerals and ambient waters" (Grossman 2012, 39). This is a thermodynamic effect (Urey et al. 1951) favouring a slightly higher proportion of the heavier isotope in the mineral than in the water but weakening with increasing temperature. Hence, under equilibrium conditions, the ¹⁸O/¹⁶O ratio of precipitating carbonate and phosphate minerals depends only on the temperature and on the ¹⁸O/¹⁶O ratio of the seawater; for a given ambient ¹⁸O level, lower δ^{18} O values in fossil shells imply higher water temperatures, and vice versa. From first principles Urey et al. (1951) estimated the variation of δ^{18} O with temperature to be about -0.176‰ per °C for carbonate minerals. Marchitto et al. (2014) combined new core top measurements with published data to derive new δ^{18} O-temperature calibrations for three groups of benthic foraminifera, Uvigerina, Cibicidoides and *Planulina*. They conclude (p.9):

The most extensive set of observations comes from the combination of *Cibicidoides* and *Planulina*, which exhibit a quadratic temperature dependence ranging from -0.25% per °C in cold waters to -0.19% per °C in warm waters (Eq. (9)), or -0.22% per °C over all temperatures if fit with a straight line (Eq. (8)).

Oxygen isotope paleothermometry of Paleozoic and Mesozoic marine fossils is based largely on the shells of brachiopods ("lamp shells") and bivalve molluscs, belemnite guards (the bullet-shaped ends of belemnite fossils) and conodonts (Grossman 2012; Veizer and Prokoph 2015). In the Cenozoic the tests (skeletons) of both planktonic and benthic foraminifera make a major contribution to the δ^{18} O database (Veizer and Prokoph 2015); of these, the benthic foraminiferal record is the most important for reconstructions of "bulk" or interior ocean temperatures (Cramer et al. 2009; Mudelsee et al. 2014; Zachos et al. 2001), while planktonic for aminiferal data is more relevant for reconstructing sea-surface temperatures (O'Brien et al. 2017; Pearson 2012). Calcium carbonate fossils are the most important for scientific study; all the organisms mentioned in this paragraph produce calcium carbonate fossils with the exception of conodonts, which are made of calcium phosphate.

Oxygen isotope fractionation also occurs when water evaporates. $H_2^{18}O$ molecules, which are heavier than the much more abundant $H_2^{16}O$

molecules, preferentially remain in the liquid phase, the resulting vapour being correspondingly depleted in H₂¹⁸O. Consequently rain or snow resulting from condensation of atmospheric water vapour has a lower δ^{18} O value than the original seawater. Condensation also causes fractionation, such that water vapour remaining in the atmosphere becomes further depleted in ¹⁸O. Most heat and mass transfer from the oceans to the atmosphere occurs at low latitudes, followed by transfer in the opposite direction at higher latitudes. Hence atmospheric water vapour becomes progressively depleted in H₂¹⁸O as it moves poleward such that the δ^{18} O value in precipitation decreases as the air cools. Furthermore orogenic (high-altitude) precipitation becomes progressively depleted in $\mathrm{H_{\scriptscriptstyle 2}{}^{18}O}$ as the air rises and cools; these processes are known as Rayleigh distillation; see fig. 2. Consequently snow, and ice sheets formed from it, exhibit lower δ^{18} O values at lower condensation temperatures (Oard 2005; Robin 1977). Despite a number of complicating factors, this effect is exploited as a paleothermometer in ice core studies (Oard 2005; Vardiman 1996a). Ice sheets today typically show δ^{18} O values between -30% and -50% (Oard 2005; Pearson 2012; Ravelo and Hillaire-Marcel 2007).

Since $H_2^{18}O$ is depleted in ice, the greater the global volume of ice, the greater the inventory of $H_2^{18}O$ residing in the oceans and atmosphere. Thus when there are significant ice sheets on the earth (as in the Pleistocene or as in Antarctica and



Fig. 2. Schematic of δ^{18} O cycle of water vapour from the ocean to the Antarctic ice sheet. Figures in the ocean and atmosphere are δ^{18} O values in ‰ (parts per thousand relative to SMOW). From Robin (1977), Fig. 7.

Greenland today), the δ^{18} O value in seawater and hence in marine fossils becomes a proxy for total ice volume: high values of δ^{18} O in the oceans imply large ice volumes. This effect was first established by Shackleton (1967) and then further developed by Hays, Imbrie, and Shackleton (1976), who linked the Croll-Milankovitch astronomical theory of Pleistocene glacial intervals ("ice ages") to the fluctuations in $\delta^{18}O$ values found in ocean floor sediments. The astronomical theory proposes that variations in the earth's orbital parameters, specifically eccentricity, obliquity (axial tilt) and axial precession, lead to variations in incoming solar radiation in summer at high latitudes (usually 65°N); glacial intervals are supposedly triggered when this radiation becomes too weak (Hays, Imbrie, and Shackleton 1976; Zachos et al. 2001); hence the title of the well-known and muchcited Hays, Imbrie, and Shackleton (1976) paper, viz. "Variations in the Earth's Orbit: Pacemaker of the Ice Ages."

Because of the apparent explanatory success of the astronomical theory, it now serves as a ruling paradigm in climate research and is often used to calibrate climate data chronologically, notably isotope records in deep-sea cores. However, this theory suffers from serious methodological flaws, described in detail by Oard (2005, 2007) and by Hebert (2016a, 2016b, 2016c). In summary, assuming the uniformitarian chronology, the most critical problems are: (1) orbitally-induced changes in solar radiation are very weak, especially the 100ka eccentricity cycle which appears, within the "astronomical forcing" paradigm, to have dominated climate variations over the last 400ka; (2) ice age cycles are in phase between Northern and Southern hemispheres, whereas the astronomical theory predicts the opposite-when the Northern hemisphere cools, the Southern should warm up, and vice versa; (3) causality, i.e. the warming which terminates a glacial interval sometimes precedes by several thousand years the change in solar radiation supposedly causing it; (4) the assumed age of 700ka for the Brunhes-Matuyama (B-M) magnetic reversal boundary, used in constructing age models for the analysis of Hays, Imbrie, and Shackleton (1976), has since been revised significantly upward to 780 ka (Hebert 2016c); (5) Hays, Imbrie, and Shackleton (1976) based their climate reconstructions on data from planktonic rather than benthic foraminifera. Since the former are much more likely to be affected by local short-term temperature variations (Hebert 2016a), how could Hays and co-authors be sure that their measured δ^{18} O values represent a globally synchronous signal? Oard (2007) claims that the uniformitarian paradigm in this field continues as a "reinforcement syndrome," and Hebert (2016a)

concludes that the data sets have "evolved" over time, thus calling into question their supposedly objective nature.

When significant ice sheets exist on earth, other climate indicators are needed to quantify separately the contributions of temperature and ice volume to δ^{18} O values, notably paleontological and geological evidence of lower sea levels and of the existence and extent of ice sheets (e.g. Clark et al. 2009; Francis 1988; Hambrey, Ehrmann, and Larsen 1991; Zachos et al. 2001). In a Flood scenario rapid changes in ocean conditions are expected, which may invalidate some conclusions based on the usual uniformitarian assumption of temporal near-equilibrium of oceanic δ^{18} O levels; detailed consideration of this issue is beyond the scope of this article.

As anticipated by Urey et al. (1951), several factors complicate the use of fossil δ^{18} O as a proxy for seawater temperature. These are summarized by Grossman (2012) as: disequilibrium precipitation of biogenic carbonate; the constancy of seawater δ^{18} O; ecological influences; spatial variability versus temporal trends; and the preservation of the record through geological time. Here the major sources of uncertainty and complexity will be considered under the headings of (1) preservation of the original fossil δ^{18} O; (2) "vital" or biological effects; and (3) environmental effects, notably seawater ¹⁸O levels. There is necessarily some interaction between environmental and vital influences because marine organisms usually respond in some way to changes in their environment (e.g. Hahn et al. 2012).

Preservation

The issue is whether oxygen isotope ratios in the fossil shell material are maintained intact or modified by the higher temperatures and pressures of burial during diagenesis or by subsequent chemical reactions. Oard (2003) cites analysis by Schrag (1999) of oxygen isotope records of planktonic foraminifera apparently indicating that Late Eocene and Oligocene tropical sea surface temperatures were up to 8°C lower than today's, yet this part of geological history was supposedly characterized by a generally warm climate. Schrag's solution was to invoke alteration by diagenetic recrystallization, which involves incorporation of relatively ¹⁸O-rich bottom water into the shells. As electron microscopy has shown, calcite can be added inside the shell or even replace its original structure, both effects being undetectable by standard microscopy (Oard 2003). Pearson et al. (2001) analyzed planktonic foraminifer shells from impermeable, clay-rich Late Cretaceous to Eocene deposits in Tanzania. Their samples were examined by electron microscopy to ensure that they were "exceptionally well preserved." The

inferred sea surface temperatures were 28-32°C, considerably higher than the figures given in most comparable previous studies. This further implies that recrystallization could have introduced major errors in previous studies, especially for planktonic foraminifera; fossils of benthic organisms such as brachiopods, bivalve molluscs, and benthic foraminifera are less prone to recrystallization because bottom water is consistently cold (typically ~2°C). Poor preservation of shell material may be indicated by opague rather than translucent shell material as noted by Grossman (2012) for brachiopods and by Wycech, Kelly, and Marcott (2016) for planktonic foraminifera. Grossman (2012) emphasizes the need for extremely careful sample screening against chemical alteration before undertaking oxygen isotope paleothermometry and describes very detailed screening protocols.

Pérez-Huerta, Coronado, and Hegna (2018) systematically review the role of biomineralization in our understanding of the fossil record. They first note that biominerals consist of a mineral phase and a multicomponent organic matrix in varying proportions, although fossils often retain little or none of the organic phase. They then systematically describe: (1) the basic characteristics of biominerals. Although there is an immense variety of these in existence, they share a number of key recognizable viz. hierarchical organization. characteristics. biocomposite nature, unique mineralization mechanisms, biological crystallographic control, and common nanostructure organization (Pérez-Huerta, Coronado, and Hegna 2018, section 2); (2) how to recognize primary (unaltered) biominerals in fossils. Preservation of the original mineralogy depends mainly on whether it has crossed its solubility threshold, i.e. whether it has suffered dissolution because of unfavorable pressures, temperatures, or water chemistry arising during its diagenetic history. Thus for calcium carbonate fossils the original mineral form, e.g. aragonite or calcite, is important because aragonite is more soluble than low-Mg calcite; (3) the factors involved in diagenetic alteration of fossils and the techniques available for evaluating such alteration.

High-resolution modern techniques including scanning electron microscopy, energy-dispersive cathodoluminescence spectroscopy. microscopy, electron backscatter diffraction and Raman spectroscopy are now used to characterize the primary mineralogy of fossils and to detect and evaluate the effects of diagenetic alteration. These have recently been supplemented by atomic force microscopy (AFM) and field emission secondary electron microscopy (FEG-SEM) which can show biomineral structures down to nanoscale. A good example of the use of such techniques is the discovery by Balthasar et al. (2011) of relic aragonite in brachiopod shells from Ordovician and Silurian rocks; detailed characterization of the shell mineralogy was key to this discovery. However an important corollary is that in earlier studies of environmental indicators in fossils, undertaken before these highly sophisticated techniques became available, the exact state of preservation of the fossils was unknown, implying that the reliability of the results, e.g. temperature reconstructions from δ^{18} O measurements, is uncertain.

Vital effects

Marine organisms which produce fossils of interest for temperature reconstructions generally live at the same temperature as the ambient water (Urey et al. 1951). However biomineralization, the formation of their hard parts, is biologically controlled and therefore does not necessarily occur in isotopic equilibrium with seawater. Hence δ^{18} O values in biominerals may not be truly representative of ambient conditions.

Examples of identifiable vital effects in planktonic foraminifera were reported by Bemis et al. (1998) in laboratory culture experiments on the symbiotic species Orbulina universa and the non-symbiotic Globigerina bulloides; the temperature range was 15–25°C. O. universa, which lives in the photic zone, hosts photosynthetic algal symbionts which are thought to modify the carbonate ion $[CO_2^2]$ concentration and hence pH locally depending on the light level (Pearson 2012). This in turn modifies the $\delta^{18}O$ difference between the shell calcite and the seawater ($\delta^{18}O_{sc} - \delta^{18}O_{sw}$) such that in high light (~16 times stronger than low light1) and with normal ambient CO_3^{2} levels, $\delta^{18}\text{O}_{sc}$ is depleted on average by 0.33‰ relative to its low-light value. This change corresponds to temperatures overestimated by $\sim 1.5^{\circ}$ C, possibly explaining a widely observed discrepancy of this magnitude between temperatures based on standard T/618O calibrations and temperatures measured in-situ for planktonic and shallowwater benthic foraminifera (Pearson 2012). In the chambered non-symbiotic G. bulloides, the main vital effect is ontogeny, which means that during shell growth successive chambers are progressively enriched in ¹⁸O relative to earlier ones; this introduces a shelleffect into temperature– $(\delta^{18}O_m)$ – $\delta^{18}O$) size calibrations. Thus for example the slope of the $T/\delta^{18}O$ curve for an 11-chambered shell is $-5.07(\pm 0.22)$, and for a 13-chambered shell $-4.77(\pm 0.27)$.

Fontanier et al. (2006) studied microhabitat and seasonal effects on the stable oxygen and carbon

isotopes in seven species of benthic foraminifera. They found that while Uvigerina peregrina forms its test in close equilibrium with bottom water δ^{18} O, all the other species (including U. mediterranea, a sister species within the same genus) maintain different but practically constant offsets against calcite formed in equilibrium with bottom water δ^{18} O. They found no systematic relationship between the foraminiferal habitat depth within the bottom sediment and the foraminiferal $\delta^{18}O$ offset against calcite in equilibrium with bottom water (designated $\Delta \delta^{18} O \equiv \delta^{18} O_{sc} - \delta^{18} O_{sc}$). Basak et al. (2009) undertook a similar isotopic study of five species of live benthic foraminifera. They found little interspecific variation in $\Delta\delta^{18}$ O values, and no systematic variation with habitat depth, which they understood to imply that metabolic rate, which depends on nutrient supply and oxygen level, does not influence ¹⁸O fractionation in these organisms.

Ishimura et al. (2012) investigated the isotope characteristics of multiple species of benthic foraminifera and found not only interspecific variations in $\Delta \delta^{18}$ O but also intraspecific variations (i.e. between individuals of the same species). In all species investigated they found an ontogenetic effect: ¹⁸O enrichment in the shell calcite increased with specimen weight, the largest specimens having $\delta^{18}O_{a}$ values closest to equilibrium calcite. Ishimura et al. (2012) concluded that (1) the species showing least intraspecific variation in $\Delta \delta^{18}$ O are the most suitable for temperature reconstructions, their best candidate being Bulimina aculeata, and (2) carbonate ion $[CO_3^2]$ concentration may be a major contributory factor to the $\Delta \delta^{18}$ O variations seen in the most variable species. Bhaumik et al. (2017) investigated microhabitat-related isotope variations between pairs of benthic foraminiferal species with overlapping ranges of habitat depth and found that only Bulimina marginata showed δ^{18} O enrichment with increasing habitat depth, interpreted as due to lower carbonate ion concentrations and hence reduced pH. Other possible vital effects noted but not directly investigated by Bhaumik et al. (2017) include respiration and "kinetic effect"; the latter, or kinetic fractionation, occurs in unidirectional (non-equilibrium) reactions and enhances the resulting degree of fractionation. A good example is photosynthesis, in which plants preferentially absorb ¹²C from atmospheric CO₂ and therefore become depleted in ¹³C relative to inorganic carbon.

In order to construct a consistent isotope record, isotope fractionation effects and different life styles

¹ High light is defined by Bemis et al. (1998) as >380 μ Einst m²s⁻¹, low light as 20–30 μ Einst m²s⁻¹. These units (which are nonstandard and no longer used) refer to the flux of *photosynthetically active photons*, i.e. with wavelengths in the range 400–700 nm. The quoted figures probably (roughly) correspond to bright sunlight and dark overcast conditions respectively.

of the organisms must be calibrated to a common standard. The most useful organisms for this purpose are brachiopods, which are represented by fossils throughout the Phanerozoic and are distributed across practically all benthic marine environments (Bitner and Cohen 2013; Wright 2014). Hence Prokoph, Shields, and Veizer (2008) introduce an "articulate brachiopod standard" (ABS), noting that: (1) the articulate brachiopod group has a stratigraphic range from the Cambrian to the present; (2) brachiopod habitats and their vital effects have been studied; (3) the diagenetic alteration of the shells can be evaluated; (4) brachiopod isotope data are available for almost the entire Phanerozoic; and (5) their primary shell material consists of low-Mg calcite for which a transfer function to water temperature can in most cases be established.

The question of how vital effects influence proxy measurements is reviewed for bivalve molluscs, cephalopods and brachiopods by Immenhauser et al. (2016), who consider traditional proxies including carbonate δ^{18} O, δ^{13} C, Mg/Ca, and Sr/Ca values, and more novel proxies including carbonate clumped isotopes (designated Δ_{47} , discussed in Part 2 of this series). Immenhauser et al. (2016) conclude with regard to δ^{18} O data from molluscs and brachiopods (p.29) that:

Under constant temperature and a non-stressed environment most molluscs and brachiopods fractionate oxygen isotopes near equilibrium with the ambient aquatic medium, at least with reference to some portions of their exoskeletons or endoskeletons and when corrected for relevant parameters.

However Immenhauser et al. (2016) also state that unanswered questions remain as to the impact of organic matrices and metastable precursor carbonate on oxygen isotope fractionation. They note that, although brachiopods generally secrete their primary (inner) calcite fibers close to oxygen isotopic equilibrium with the seawater, some studies have reported kinetic and metabolic isotope fractionation differing between taxa and even within a single brachiopod valve, and that from the bulk isotopic range across the secondary shells of two congeneric brachiopod species seawater temperatures were overestimated by 24°C. Much of this kind of variation is now thought to be due to a "MgCO₃ effect" (Brand et al. 2013); incorporating MgCO₂ into brachiopod shell calcite affects its $\delta^{18}O$ level and hence the T/ $\delta^{18}O$ calibration equation. Having investigated this effect in several species of modern brachiopods from a range of globally-distributed sites, Brand et al. (2013) proposed a new calibration equation for articulated brachiopods:

 $T(^{\circ}C) = 16.192 - 3.468(\delta^{18}O_{sc} - \delta^{18}O_{sw} - \delta_{M\sigma}) \qquad (2)$

where $\delta^{18} O_{sc}$ and $\delta^{18} O_{sw}$ are the shell calcite and seawater $\delta^{18} O$ levels respectively and δ_{Mg} is an adjustment in $\delta^{18} O$ of +0.17‰ per mol% MgCO₃ incorporated into the shell calcite. Although $\delta^{18} O_{sc}$ and δ_{Mg} can in general be measured for a single fossil shell, use of equation (2) presupposes that $\delta^{18} O_{sw}$ is either known or can be assumed from additional information. Brand et al. (2013) suggest that similar adjustments for the MgCO₃ content of calcite in inarticulated brachiopods and other high-Mg calcitic marine invertebrates would enable their use as valuable paleotemperature archives.

It is perhaps worth noting here that while biomineralization, specifically the formation of calcitic shells by marine invertebrates, is biologically controlled it is also affected by seawater chemistry (Pérez-Huerta, Coronado, and Hegna 2018). Thus for example calcitic bivalve molluscs grown in artificially high-Mg seawater were found to secrete aragonite on their interior shell surfaces (Checa et al. 2007). Variations in Mg/Ca in the oceans through the Phanerozoic and its effects will be discussed in Part 2 of this series.

Seawater ¹⁸O levels

The ¹⁸O level in fossil shell material depends on both its formation temperature and the $\delta^{\rm 18}O$ value of the seawater: we assume either that the global ice volume is negligible or that its effect on seawater δ^{18} O values has already been accounted for. There is naturally some spatial variation in the ¹⁸O level in the oceans. Thus according to Grossman (2012), the bulk of modern seawater, represented by deep water masses, has a relatively narrow δ^{18} O range from about -0.6‰ (Antarctic Bottom Water) to +0.1‰ (North Atlantic Deep Water). Open-ocean surface waters are more variable, ranging from about -0.5%in the Southern Ocean to +1.4% in the dry subtropical zone in the North Atlantic. ¹⁸O variability in the surface waters of enclosed seas, such as the Arctic Ocean, Mediterranean Sea, and Red Sea, is even larger, roughly -2% to +2%, owing to the varying balance of evaporation and precipitation. There is an acknowledged correction for the effects of latitudinal variation in the precipitation-evaporation balance in the Southern Hemisphere (Grossman 2012). Close to continents, mixing with fresh water can lower $\delta^{18}O$ by an amount depending on the δ^{18} O of the river water, which contributes to the correlation between salinity and seawater $\delta^{18}O$ (LeGrande and Schmidt 2006; Polyak, Stanovoy, and Lubinski 2003); in some circumstances this correlation means that δ^{18} O becomes primarily a proxy for water density rather than temperature (Lynch-Stieglitz, Curry, and Slowey 1999). Other factors noted by Grossman (2012) include water depth, global ice volume, which affects salinity because freezing seawater excludes salt, and precipitation, which reduces seawater δ^{18} O values locally (LeGrande and Schmidt 2006).

Spatial variability of δ^{18} O in Holocene oceans is thus reasonably well understood. However temporal variability has been controversial and harder to pin down. Practically all investigators, notwithstanding scatter and variations in both directions, have found that fossil ¹⁸O levels increase progressively upwards through the geological record. Some have treated this as a diagenetic artifact (e.g. Killingley 1983), while others have deduced a rising level of ¹⁸O in the oceans, and yet others have inferred long-term cooling. Veizer and Prokoph (2015) argue that the oceanic ¹⁸O level has risen progressively through geological time, citing: (1) the across-the-board consistency of the secular trend in ¹⁸O measurements, including those where sample integrity was carefully checked, and (2) the corresponding secular trend, confirmed independently by different groups, in other isotopic and element ratios. These include carbon, sulfur, calcium, radiogenic and stable strontium isotope ratios, and Sr/Ca ratios (Prokoph, Shields, and Veizer 2008). Veizer and Prokoph (2015, 93) state:

It is therefore inconceivable to argue that the atoms of oxygen, the dominant structural unit of calcite crystals, must have been massively exchanged during diagenesis while at the same time none of the other major and trace elements or isotopes were affected.

Hence Veizer and Prokoph conclude that the secular trend in fossil δ^{18} O values indicates increasing oceanic ¹⁸O levels rather than changing temperatures. This must have involved oxygen exchange between the water and hot silicate rocks, for example at mid-ocean spreading ridges. Interaction with rocks at temperatures above 350°C causes ¹⁸O enrichment of seawater; conversely ¹⁸O depletion occurs at lower temperatures (Walker and Lohmann 1989). Changes in seawater chemistry revealed in

coupled variations in the mineralogies of marine aragonite limestones and potash evaporites through the Phanerozoic have been attributed to variations in the midocean ridge hydrothermal brine flux, due in turn to variations in the rate of ocean crust production (Hardie 1996; Stanley, Ries, and Hardie 2002). Thus seawater chemistry, defined in terms of ion content (e.g. Mg/Ca and Sr/Ca) and isotope ratios (e.g. δ^{18} O), has undoubtedly changed significantly over geological time because of crustal movements and changes in heat and mass transfer at the ocean floor. The challenge is to integrate the changes apparent in the geological record into a Flood-based geological paradigm.

2. Appraisal of Oxygen Isotope Data

For Flood models such as CPT (Austin et al. 1994), the Cenozoic represents the post-Flood period, characterized in its late stages by the growth of ice sheets and an Ice Age (Oard 1990). As noted earlier, in other Flood models the correspondence of ocean floor development with Flood chronology will vary depending on where the Flood/post-Flood boundary is placed in the geological record; this question is considered further in section 3.

Cenozoic oxygen isotope data are reviewed by Zachos et al. (2001), based on δ^{18} O measurements on the calcareous fossil remains of two types of common, long-lived benthic foraminifera, Cibicidoides and Nuttallides. Zachos et al. treat the $\delta^{18}O$ record as a proxy for ocean floor temperature up to the late Eccene, when the Antarctic ice sheet began to form, around 35MaBP (before present) in the uniformitarian chronology. From the early Oligocene they treat $\delta^{18}O$ as a gualitative indicator of ice volume. Although the trend is not monotonic, the overall change in δ^{18} O is +5.4‰ over the whole Cenozoic, which they resolve into +3.1‰ to represent deep-sea cooling, +1.2‰ for growth of the Antarctic ice sheet and +1.1‰ for subsequent growth of Northern Hemisphere ice sheets. A more comprehensive $\delta^{18}O$ data compilation is given by Cramer et al. (2009). The highest global ocean-floor temperatures in the Cenozoic inferred from $\delta^{18}O$ data are approximately 12–13°C (see fig. 3), occurring at the Paleocene-Eocene Thermal Maximum (PETM) at about 55 Ma and during the Early Eocene Climate Optimum (EECO) at about 51 Ma in the uniformitarian chronology (Cramer et al. 2009; Mudelsee et al. 2014; Zachos et al. 2001). This temperature limit can be taken as robust since it is based on low-Mg calcite fossils with minimal MgCO₂ effect on their δ^{18} O values (cf. equation 2), and the contribution of the secular trend in seawater ¹⁸O levels identified by Veizer and Prokoph (2015) is insignificant for the Cenozoic.



Fig. 3. Global Cenozoic δ^{18} O values from benthic foraminiferal shells found in marine sediment cores. The temperature scale, which assumes an ice-free ocean, employs the transfer function used by Zachos et al. (2001). Since ice sheets appeared around the Eocene-Oligocene transition, indicated temperatures thereafter are too low, rising δ^{18} O values indicating growing ice volume as well as low temperatures. From Mudelsee et al. (2014), Fig. 1.

The review of δ^{18} O data covering most of the Phanerozoic up to the present (i.e. from 512Ma in the uniformitarian chronology) by Veizer and Prokoph (2015) includes the data published by Zachos et al. (2001) and Cramer et al. (2009), together with data from several other sources. Altogether, measurements from over 58,000 low-Mg calcite fossil marine shells are included. The Paleozoic portion of the record is covered almost entirely by brachiopods while other taxa dominate the Mesozoic and Cenozoic. Fig. 4 [taken from Fig. 1 in Veizer and Prokoph (2015)] shows, despite distinct fluctuations, a secular trend of +6±1‰ over the whole dataset. Shells secreted at low temperatures are extremely scarce; almost none from the Paleozoic plot above the 0‰ level of δ^{18} O for modern calcite.

If the observed secular trend is interpreted through a single δ^{18} O/ocean temperature calibration with present-day SMOW (Standard Mean Ocean Water) as the 0‰ reference, it implies considerable cooling from 50–55°C in the mid-Cambrian to a presentday average of 15°C; ocean temperatures today lie between about 0°C and a sharply-defined maximum of 30–32°C on the surface (NOAA 2018). As noted previously, Veizer and Prokoph argue rather that the δ^{18} O trend is due to a secular increase in the ¹⁸O level in ocean water since the Cambrian.

Oxygen-18 levels in Precambrian strata show a similar trend (Jaffrés, Shields, and Wallmann 2007; Prokoph, Shields, and Veizer 2008; Shields and Veizer 2002; Veizer and Hoefs 1976), though in uniformitarian terms the change is slower than in the Phanerozoic. Thus Precambrian δ^{18} O values are lower than Phanerozoic values. Some early papers inferred very high Precambrian water temperatures. For example, Knauth and Epstein (1976) deduced from δ^{18} O values in Archean cherts that water temperatures were up to 70°C at 3Ga (billion years) ago, but others have questioned the accuracy of this parameter as a temperature proxy. Thus Walker (1982) contends that, since evaporitic gypsum cannot form above 58°C vet is found in Precambrian strata, the water temperature could not have been above 58°C. Walker also sees the continued presence of life in the Precambrian as implying temperatures not exceeding 60°C. In contrast, Knauth and Lowe (2003) and Knauth (2005) argue from $\delta^{18}O$ measurements of cherts from the Swaziland Supergroup that Archean temperatures lay in the range 55–85°C. Further studies suggest Archean ocean temperatures much closer to present-day values. Thus from analysis of both oxygen and hydrogen isotope ratios in an Archean chert, Hren, Tice, and Chamberlain (2009) deduce ocean temperatures not exceeding 40°C, while temperatures derived from oxygen isotope measurements in Archean phosphates lie in the range 26-35°C (Blake, Chang, and Lepland 2010), roughly equivalent to present-day values. A recent study of volcanic rocks and cherts from the Barberton Greenstone Belt in South Africa (De Wit and Furnes 2016) reveals the existence of deep-water



Fig. 4. δ^{18} O of Phanerozoic low-Mg fossils (57,146 data points), taken from Fig. 1 in Veizer and Prokoph (2015). Temperature estimates are based on the Visser, Thunell, and Stott (2003) transfer function, assuming the presentday δ^{18} O value of 0‰ SMOW for seawater. Veizer and Prokoph (2015) give details of all the data and their sources.

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hydrothermal systems in a relatively cold Archean environment; temperatures inferred from $\delta^{18}O$ data range from 30°C to above 200°C. Could part of the reason for apparently conflicting temperature indications from studies of Archean rocks be an extremely heterogeneous formation environment in terms of temperature and water chemistry?

3. Interpretation of the Data in Terms of Flood Geology

All discussions of complicating issues in the secular literature presuppose the uniformitarian chronology, within which the assumption of a climate system approximating to steady state is a meaningful concept. Inarticulated brachiopods typically live for \sim 1–10 years and articulated brachiopods for \sim 3–30 years (Bitner and Cohen 2013; Emig 1997; Immenhauser et al. 2016), while bivalves can live for 50 years or more and deep-water species up to 100 years (Jones 1989). These life spans are short compared with climate change time scales typically contemplated within uniformitarian geology; in that framework the concept of quasi-steady state may be appropriate.

However, within a biblical timescale, in which (say) most of the Phanerozoic represents deposits laid down during the Flood year, conditions must have changed rapidly. Thus in a CPT Flood scenario, there will probably have been rapid, large-scale exposure of hot mantle rock to ocean water, resulting in considerable exchange of oxygen and significant change in the oceanic ¹⁸O level. Furthermore the breaking up of "the fountains of the great deep" (Genesis 7:11), perhaps as a result of the dewatering of rising, depressurizing magma, could have injected into the oceans a large quantity of water with a different δ^{18} O value from that of the pre-Flood ocean. Unless it has suffered recrystallization, the calcite in fossil shells deposited during the Flood year, many of which represent animals killed and buried in the Flood, must represent ambient conditions immediately before or in the early stages of the Flood. If ¹⁸O levels changed significantly during the Flood (though the direction and magnitude of change cannot be predicted), this change should be reflected in the calcite in post-Flood fossil shells. In the standard CPT Flood model (Austin et al. 1994) these would be the lowest-level Cenozoic fossils, i.e. from the Paleogene; in Flood models which place the Flood/ post-Flood boundary higher in the stratigraphic record these fossils would be found correspondingly higher. Inspection of fig. 3 suggests no obvious major change in the δ^{18} O level either at the K-Pg interface or above it except possibly at the PETM or EECO.

Some who accept CPT as the basic explanatory framework for the Flood would now place the Flood/ post-Flood boundary somewhere between the K-Pg boundary and the Pliocene/Pleistocene boundary. Thus Baumgardner, who developed a coherent physical model for the mechanism underlying CPT (Baumgardner 1986, 1994a, 1994b, 2003) and was a co-author of the Austin et al. (1994) paper, places the Flood/post-Flood boundary in the Pliocene (Baumgardner 2005, 2012). Snelling (2009) argues that this boundary must lie somewhat above the K-Pg interface on account of the large amount of geological work in evidence in rocks above the Cretaceous. Clarey (2016) firmly supports the CPT model but does not say exactly where he would place the Flood/post-Flood boundary, while Ross (2012), who also subscribes to CPT, inclines to the standard CPT placement of this boundary.

Not all Flood models view the geological record in this way. Brand (2007) assumes considerable geological activity both before and after the Flood, though he sees the Flood as the single most significant event in earth history. His model does not predict a clear end-of-Flood marker in the geological record. Gentet (2000), in his CCC (Creation/Curse/ Catastrophe) model, acknowledges the order of the stratigraphic column as real and treats the different fossil assemblages in the geological record as representing different ecological zones rather than geological periods. Consequently, although Gentet believes that Precambrian rocks date from Creation Week, he places the pre-Flood/Flood and the Flood/post-Flood boundaries in different strata within the geological column in different locations; however most creation scientists view the column as lithostratigraphically and biostratigraphically real whilst maintaining caution about the details (Snelling et al. 1996; Tyler and Coffin 2006). Oard (2016), who has written extensively on Flood-related issues but who does not accept CPT, adopts the "Late Cenozoic Boundary Model," which places the Flood/ post-Flood boundary at different stratigraphic levels in different locations, but all corresponding to a "Late Cenozoic" designation. None of these alternatives provides a coherent physical mechanism to explain the Flood in the same way as CPT, nor is it possible to make any meaningful comparison between model predictions and data for seawater temperatures through the Flood/post-Flood boundary except where a clear end-of-Flood marker has been defined.

If the geological record is understood in terms of a Recolonization model, in which Flood deposits correspond largely to Hadean and Archean rocks (Tyler 2006), practically all accessible fossil marine shells will have been deposited over the subsequent 4400 or so years. Since these shells would not then have been deposited during the Flood, steady-state conditions might be a better approximation to reality. However, Tyler envisages that the recovery and recolonization of the earth after the violence of the Mabbul would have been punctuated by further catastrophic events, implying that for much of geological history one cannot necessarily assume steady state. Given the stratigraphically poorlydefined end to the Flood in this model, end-of-Flood seawater temperature predictions for this model cannot be compared with geological data, although Tyler (2006) suggests that the high Archean water temperatures indicated by the data of Knauth and Lowe (2003) and Knauth (2005) would be consistent with hot ocean water at the end of the Flood.

A number of creation scientists (e.g. Austin et al. 1994; Vardiman 2013) have taken at face value published paleotemperatures based on $\delta^{18}O$ data, apparently without considering the validity of the δ^{18} O/temperature calibration used in their sources. However Vardiman (1996b) outlines the basics of oxygen isotope thermometry and discusses many of its complicating factors. He cites: (1) biological factors (e.g. disequilibrium effects in foraminifera); (2) diagenetic effects (e.g. recrystallization, biasing of a fossil assemblage due to preferential solution of the "warmer" fraction); (3) ecological factors (e.g. seasonal migration), which must be accounted for in linking inferred temperatures with climate: (4) in a Pleistocene context, foram habitat changes between glacial and interglacial conditions. Vardiman (1996b) also acknowledges the impact of the seawater $\delta^{18}O$ value on temperature estimates, but only in relation to the growth and retreat of Pleistocene ice sheets. Vardiman does not mention the effect on ¹⁸O levels of interactions between hot mantle rocks and seawater, nor does he consider the possibility of a long-term trend in global oceanic ¹⁸O levels. He concludes that in a Flood model context the accuracy of past temperature estimates is unlikely to be better than 3 or 4°C.

Regardless of the particular Flood scenario assumed, creation scientists must inevitably interpret the ¹⁸O record in marine fossils differently from uniformitarian scientists, mainly because their timescales are much shorter. Furthermore creation scientists have not thus far considered the possibility (discussed above) of a long-term increase in ambient ¹⁸O levels. Thus further investigation is needed to clarify our interpretations of δ^{18} O values in pre-Cenozoic marine fossils and their implications for past ocean temperatures.

Surface Heat Flows 1. General Picture

The earth is steadily losing internal heat. According to Pollack, Hurter, and Johnson (1993), the total outward surface heat flow is 44.2×10^{12} W, or 44.2 TW, which implies an average heat flux of 87 mWm^{-2} . Davies and Davies (2010) estimate a figure of (47 ± 2) TW for the total outward heat flow; the uncertainty of ± 2 TW is their 20 statistical uncertainty. The division between continents and oceans given by Davies and Davies is shown in Table 1.

Table 1. Partitioning of surface heat flows between continents and oceans. After Davies and Davies (2010).

	Area (km²)	Heat Flow (TW)	Mean Heat Flux (mWm ⁻²)	
Continent	2.073×10 ⁸	14.7	70.9	
Ocean	3.028 × 10 ⁸	31.9	105.4	
Global total	5.101 × 10 ⁸	46.7	91.6	

The major contributions to the total surface heat flow are radiogenic heat production in the crust and mantle, heat flow from the core, and secular cooling of the mantle (Furlong and Chapman 2013). These authors estimate that approximately half the earth's total heat budget is due to heat released from the core through the CMB and to radiogenic heat generated in the lower mantle. Partitioning of the remainder (20–25TW) between radiogenic heat production (largely in the continental crust) and secular cooling of the mantle is the subject of debate among uniformitarian scientists.

According to Furlong and Chapman (2013). continental heat flow largely arises from radiogenic heat production in the lithosphere (notably the continental crust), while oceanic heat flow is dominated by lithospheric cooling away from mid-ocean ridges (MORs). Fig. 5 (their Figure 1) schematically illustrates the partitioning of surface heat flows. The temperature field in the oceanic lithosphere is well constrained, the surface heat flux declining with distance from the MOR. In contrast, the continental temperature field is complex and generally poorly resolved because vertical heat generation profiles are not well known, there is considerable lateral variation in heat production and continental rocks have undergone a complex tectonic history.

Below the plate boundary corresponding to a mid-ocean ridge heat is transported upwards by upwelling hot mantle material participating in mantle convection. As this material rises and phase changes due to decompression take place, the axial heat flux consists of (1) the release of latent heat of crystallization of basaltic magma, and (2) the heat lost by cooling of magma down to hydrothermal temperatures \sim 350°C (Elderfield and Schultz 1996). The resulting rock accretes to the base of the spreading plates, leading to plate thickening with distance from the MOR. According to Turcotte and



Fig. 5. Schematic of heat flow partitioning at various levels in both oceanic and continental lithosphere. From Furlong and Chapman (2013), Fig. 1.

Schubert (2002, section 4.16) the lower boundary of the plate may conveniently be defined by the 1600 K isotherm, which corresponds to the temperature below which mantle rocks do not readily deform on conventional geological timescales; however the exact choice of isotherm temperature is not critical for subsequent analysis. Turcotte and Schubert (2002, section 4.16) ignore hydrothermal effects and undertake a formal heat conduction analysis which gives a plate thickness of $\sim 2(\kappa t)^{\frac{14}{5}}$, where κ is the thermal diffusivity of the plate material (taken to be $10^{-6} \text{ m}^2 \text{s}^{-1}$) and t is the time since moving away from the MOR. Fig. 5 qualitatively depicts the plate thickening process.

However Furlong and Chapman (2013) view the rheologically plate-like behavior of oceanic lithosphere as resulting from its compositional properties, viz. water content, rather than simply from its thermal properties, and suggest that this has important consequences for the transfer of heat from the deeper mantle into the lithosphere. They first note that mantle material is emplaced at shallow levels as a result of partial melting and separates into crustal and residual mantle components (see fig. 5). Furlong and Chapman cite the work of Hirth and Kohlstedt (1996) who reevaluated high-pressure experimental data on the viscosity of olivine aggregates (dunite) in relation to water content, arguing that olivine is the rheologically dominant component of peridotites. Hirth and Kohlstedt conclude that there is an excess of water in the MOR basalt (MORB) source region, corresponding to a pressure of 300MPa, which implies that the viscosity there is hundreds of times lower than in dry olivine aggregates. Hirth and Kohlstedt thus estimate that the partial melting which generates new oceanic crust starts ~115km

down and dehydrates the associated mantle, leading to a mechanical layering of dry mantle above the 60–70km depth interval with wet mantle beneath. Consequently viscosity decreases by 2–3 orders of magnitude from above to below depths of approximately 70–100km. Hirth and Kohlstedt (1996, Abstract) conclude:

These observations indicate that the base of an oceanic plate is defined by a compositional rather than thermal boundary layer, or at least that the location of the thermal boundary layer is strongly influenced by a compositional boundary, and that the evolution of the oceanic upper mantle is strongly influenced by a viscosity structure that is controlled by the extraction of water from olivine at mid-ocean ridges.

Furlong and Chapman (2013) deduce that there is little upward conductive heat flow at 100 km depth; the conductive heat flux at the base of the plate increases to no more than about 17 mWm⁻² as the plate material cools and moves away from the MOR; for comparison, the average surface oceanic heat flux is 105.4 mWm⁻² (see Table 2), a typical minimum is 48 mWm⁻² (Stein and Stein 1992), and the maximum in MORs is of order 450 mWm⁻² (Davies and Davies 2010). Thus since very little radiogenic heat is produced within the oceanic lithosphere ($\approx 4 \,\mathrm{mWm^{-2}}$, Parsons and Sclater 1977), most heat emerging from young ocean regions must result from secular cooling. As indicated in fig. 5, the surface heat flux decreases with distance from the MOR, and the basal heat flux from the mantle below 100km, although small, becomes a progressively larger fraction of the total (Furlong and Chapman 2013). Furlong and Chapman estimate that a total of $\sim 3.9 \times 10^{14}$ Joules (per square meter of surface

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Thermally Important Radioactive Elements in the Earth						
Isotope	Energy/Atom (MeV)	µW/kg of Isotope	μW/kg of Element	Estimated total earth content (kg)	Total heat (TW)	Total heat 4.5 Ga ago (TW)
²³⁸ U	47.7	95.0	94.35	12.86×10 ¹⁶	12.21	24.5
²³⁵ U	43.9	562.0	4.05	0.0940 × 10 ¹⁶	0.53	44.4
²³² Th	40.5	26.6	26.6	47.9×10 ¹⁶	12.74	15.9
⁴⁰ K	0.71	30.0	0.0035	7.77×10 ²⁰	2.72	33.0
				Totals	28.2	117.8

Table 2. Heat outputs from thermally important radionuclides. After Stacey and Davis (2008), Table 21.2.

area of fresh oceanic lithosphere) is deposited in the upper 100 km of the region, an enormous heat load if we are considering biblically-compatible timescales.

Stacey and Davis (2008), writing in uniformitarian terms, consider both ocean depth and ocean floor heat flux as functions of age, using the data and models of Stein and Stein (1992) and Stein (1995). The first cooling model used by these authors is a simple half-space model, with heat loss at the ocean floor both by conduction and by hydrothermal circulation of seawater. As the oceanic lithosphere cools, it contracts, its density increases, and its upper surface subsides. Consequently ocean depth increases, an effect enhanced by the increased water loading. However, model plots based on this analysis do not fully reproduce the data; after 30Ma the heat flux remains roughly constant, and after 70 Ma the ocean depth levels out. A preferred alternative known as the plate model assumes that mature lithosphere reaches a limiting thickness, after which it ceases to cool further and behaves like a plate in steady state, its lower boundary maintained at a fixed temperature by contact with the asthenosphere (Richardson et al. 1995; Stein and Stein 1992). Another alternative assumes that the lower boundary of the oceanic lithosphere is at a fixed temperature (i.e. it coincides with an isotherm) and is subject to a constant upward convective heat flux; this is known as the CHABLIS (Constant Heat flow Assigned on the Bottom Lithospheric ISotherm) model (Crough 1975; Doin and Fleitout 1996). Both plate and CHABLIS models produce good fits to surface heat flux data for old lithosphere, but CHABLIS predicts continuing subsidence well beyond 80Ma, and the proportion of global mantle cooling due to erosion of material from the base of the lithosphere by secondary convection is 40% (the remaining 60% is due to subduction), much higher than in the plate model (Doin and Fleitout 1996). Although both models predict a lithospheric thickness (defined by a particular isotherm) of about 80km after 100Ma, the CHABLIS lithosphere continues to thicken well beyond this point. Another widely observed feature of oceanic heat flows is that all conduction-based models significantly overpredict surface heat fluxes for relatively young

oceanic lithosphere; this is attributed to ventilated hydrothermal circulation recharge and discharge at crustal outcrops or areas of thin sediment cover (Hasterok, Chapman, and Davis 2011).

In the plate model the asthenosphere, unless it is continuously and rapidly replaced, can only supply the heat flow needed to keep the base of the lithosphere at a fixed temperature by cooling, which must lead to lithospheric thickening. Stacey and Davis (2008) note that: (1) there is indeed seismic evidence of continuing lithospheric thickening, and (2) although hydrothermal cooling of young ocean lithosphere is certainly observed and modifies the vertical temperature profile through the upper few kilometers, attempts to model it had not been fully successful. Since surface heat flux and ocean depth stabilize on different time scales, Stacey and Davis suggest that two additional effects need to be modeled. They raise two possibilities: (1) hydrothermal circulation extending deep into the lithosphere, redistributing heat but not necessarily exhausting it into the ocean; (2) an asthenosphere which thickens as it approaches a subduction zone. However they do not consider that a fully satisfactory solution is available.

Given that the earth's ocean basins are geologically young, few areas being older than early Jurassic (Müller et al. 2008), and that most creation scientists regard Jurassic rocks as Flood deposits, these basins must have formed during and since the Flood, and most oceanic lithosphere must have cooled to its present state within that time, i.e. in no more than 4500 years, probably far less. As noted above, the heat load of $\sim 3.9 \times 10^{14}$ Jm⁻² of fresh ocean lithosphere due to material surfacing at MORs (Furlong and Chapman 2013) is enormous, more than 30 times enough to boil off the oceans if deposited very rapidly. The associated "heat problem" is to determine how the cooling was accomplished in a short time (Barnes 1980). A first attempt at this was made by Worraker and Ward (2018), who used the framework provided by the plate model of Stein and Stein (1992), suitably modified to accommodate biblically-compatible time scales, by invoking a transient subsurface heat sink. This was treated as a function of both space and time combined with rapid but decelerating plate motion. However, even with the optimum choice of model parameters available, Worraker and Ward were unable to find even a remotely satisfactory solution: predicted surface heat fluxes were far too high and ocean depth profiles far too sharp. These problems stem from the existence of a narrow near-surface thermal boundary layer, an inevitable consequence of the short time scales imposed in the model. Further details of this work and a discussion of possible ways forward in solving this heat problem, probably including hitherto neglected geophysical effects (e.g. deep hydrothermal convection, superheated steam jets, etc), will be given in Part 4 of this series.

2. A Creationist Resolution of Problematic Heat Flow Data

Baumgardner (2000) has shown that a longstanding puzzle for the earth science community, namely the widely-found correlation between surface heat flow and the measured radioactivity of surface continental igneous rocks, can be neatly explained within the creationist paradigm. Uniformitarian explanations appear rather contrived. For example, (1) Roy, Blackwell, and Birch (1968) invoke a vertical column of rock 7-11km deep through which the level of radioactivity is practically uniform and is closely correlated with the surface rock type, and (2) Lachenbruch (1970) suggests a model in which radiogenic heat production decreases exponentially with depth. Subsequent investigations in crystalline continental crust elsewhere (e.g. in Russia and in Germany) have also shown no variation of radiogenic heat production with depth-rather, it follows lithology (Clauser et al. 1997; Pribnow and Winter 1997). The problem is discussed by Furlong and Chapman (2013), who prefer Lachenbruch's (1970) exponential model to the alternatives, but do not offer a definitive solution.

Baumgardner (2000) proposed a modest pulse of accelerated nuclear decay (AND) during the Genesis Flood (~5000 years ago) which produced an episode of increased radioactive heating whose aftermath now dominates the surface heat flow in the host plutonic rocks. This would have raised the temperature of the host rocks by an amount proportional to the concentration of heat-generating elements. Baumgardner estimates that today's observed heat flows can be explained in terms of 190,000 years' worth of accelerated decay through a depth of approximately 1km. The resulting temperature rise lies between 5.5K and 22K, corresponding to the approximate lower and upper limits respectively of present-day radiogenic heating in granites. These figures are perfectly credible and the depth requirement is more plausible than the uniformitarian assumption that the surface level of radioactivity is correlated downwards to depths of about 10km.

From his contributions to the RATE project, Snelling (2005a, 2005b, 2005c, and pers. comm.) concluded that:

... the RATE project demonstrated that there was physical evidence (fission tracks and radiohalos) for only 500–600 million years' worth of accelerated radioactive decay only during the Flood.

Given that the Genesis Flood lasted about a year, Snelling's comment implies an overall acceleration factor of about 6×10^8 , or 600 million, considerably larger than the value of 1.9×10^5 implied by Baumgardner's results. To reconcile this difference note that Baumgardner's analysis deals only with the *net* additional heating of the rocks during the Flood as a result of AND; given a sufficiently strong additional cooling mechanism operating simultaneously during the Flood, there is no conflict. The challenge is to identify a cooling mechanism strong enough to limit the temperature rise in granite plutons to the ~20 K level deduced by Baumgardner.

A rough idea of the magnitude of this problem may be obtained by considering the total radiogenic heat generated during the Flood year as a result of AND within typical granite, and comparing it with the heat needed to melt the granite. Taking the average present-day radiogenic heat production in granite as 1.05×10^{-9} Wkg⁻¹ (Table 3), the total heat generated by 6×10⁸ years' worth of radioactive decay is 1.99×10⁷ Jkg¹. Taking the granite melting point as its typical liquidus temperature of 1550K, its latent heat of melting as 4.2×10^5 Jkg⁻¹ and its specific heat capacity as 830 Jkg⁻¹K⁻¹ (Stacey and Davis 2008, Table A.6), and assuming a starting temperature of 300K (=27°C), the heat needed to melt it is 1.46×10^{6} Jkg⁻¹. The predominant mineral component of granite, silica (SiO₂) has a boiling point at 1 bar of 3177K and an enthalpy of vaporization of approximately $1.2 \times 10^7 \text{Jkg}^{-1}$ (Kraus et al. 2012). Thus if delivered adiabatically (i.e. in a time short compared with the thermal diffusion time scale of the granite body, which would be the case for anything larger than ~5 m across) the AND heat load is more than 13 times enough to completely melt the granite, and may even be enough to vaporize it. Although this estimate is somewhat misleading (e.g. much granite was undoubtedly emplaced during rather than before the Flood, etc.), it suggests nevertheless that a powerful cooling mechanism operating at the same time as AND is needed. This issue is considered further in later articles in this series.

3. Inventory of Heat-Producing Radioactive Elements

Baumgardner (2000) reviews the distribution of radioactive isotopes in the earth from a creationary

		Concentration (parts per million by weight)				Heat Production
	Material	U	Th	К	K/U	(10 ⁻¹² W/kg)
Igneous rocks	granites	4.6	18	33000	7000	1050
	alkali basalts	0.75	2.5	12000	16000	180
	tholeitic basalts		0.4	1500	13600	27
	eclogites	0.035	0.15	500	14000	9.2
	peridotites, dunites	0.006	0.02	100	17000	1.5
Global mean	crust (2.8×10 ²² kg)	1.2	4.5	15500	13000	293
	mantle (4.0×10 ²⁴ kg)	0.025	0.087	70	2800	5.1
	core	0	0	29	—	0.1
	whole earth	0.022	0.081	118	5400	4.7

Table 3. Measured abundances of main heat-producing radioactive elements. After Stacey and Davis (2008), Table2.13.

perspective. He notes that mainstream models of the earth's chemical make-up are based mainly on data from present-day rocks, and are therefore generally robust and reliable. Notable features include: (1) the earth contains the same recipe of higher melting temperature elements as the sun and most meteorites; (2) it has undergone significant chemical differentiation in its history, including segregation of much of its iron to the center to form the core, and the extraction of continental crust from the silicate mantle by partial melting, which appears to have extracted and concentrated into the crust a large proportion of the incompatible elements originally in the mantle. Incompatible elements do not fit readily into the lattice structure of the dominant refractory silicate minerals because they possess either large ions or high electric charges. They include: large-ion lithophile (LIL) elements like K, Rb, Cs, U, Th, Sr, and Ba; rare-earth elements (REE) like La, Ce, Nd, Sm, and Lu; and high field strength elements (with high ionic charge) like Nb, Ta, Hf, and Ti. In particular, these lists include the earth's most important heatgenerating radioactive elements, uranium, thorium, and potassium.

Estimated heat outputs for U, Th, and K are listed in Table 2, taken from Stacey and Davis (2008, Table 21.2); the final column gives the total heat output which these isotopes would have contributed 4.5Ga ago on the uniformitarian assumption that decay constants have not changed. The energy/atom values include all series decays to final daughter products. Average locally absorbed energies are included, but not neutrino energies because neutrinos have such low interaction cross sections with most elements that practically all escape into space without depositing their energy (see below for more on neutrinos).

As noted above, heat-generating radioactive elements are not evenly distributed: they are strongly concentrated in the continental crust. Table 3 (taken from Table 21.3 in Stacey and Davis 2008) compares measured abundances in geological materials with estimated average concentrations in the main components of the earth's interior. Continental crustal rocks, notably granites, are strongly enriched in incompatible elements and hence in U, Th, and K while the mantle as a whole is correspondingly depleted, although the upper mantle contains a mixture of enriched and depleted material; hence MORBs also contain a mixture of enriched and depleted material (Gale et al. 2013). Several uniformitarian assumptions are made by Stacey and Davis (2008) and in the references they cite in order to produce their estimates. Since estimating the concentration of radiogenic elements in the (largely unobservable) mantle is inherently difficult, comparisons with other Solar System bodies (notably the moon and meteorites) and their supposed formation history are needed in order to arrive at their figures.

A measure of the uncertainty in the estimated inventories of these elements may be gleaned by comparing the figures quoted by Lodders and Fegley (1998, Table 6.8) from two reference sources with those quoted by Stacey and Davis (2008, Table 21.3); see Table 4. Although it may be argued that there must have been an overall improvement over time in the data and models used to generate these estimates (e.g. Kargel and Lewis 1993), the significant differences, by factors in the range 1.5–

Table 4. Whole earth abundances of radiogenic elements: comparison of estimates from different sources. Table entries are parts per million by weight. The sources used by Lodders and Fegley are (in their notation): MA80, Morgan and Anders (1980), KL93, Kargel and Lewis (1993).

	К	Th	U			
Stacey and Davis (2008), Table 2.13	118	0.081	0.022			
Lodders and Fegley (1998), [MA80]	135	0.0512	0.0143			
Lodders and Fegley (1998), [KL93]	225	0.0543	0.0152			

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1.9, do not inspire confidence. However, the potential heat problem due to AND identified previously would not be significantly alleviated even if, for example, the previously-assumed concentrations of U, Th, and K in granite were halved—a heat load only seven times enough to completely melt the granite is still overwhelming!

4. Radiogenic Heating Inferred from Geoneutrino Measurements

Of the earth's most important heat-producing radiogenic elements, most ⁴⁰K (89.3 %) decays into calcium by direct β^- decay, and the decay chains of both ²³⁸U and ²³²Th include several β⁻-decay steps [β⁻-decay involves the emission of electrons and antineutrinos, whereas β^+ -decay produces positrons and neutrinos]. Unlike the other products of radioactive decay (a particles, electrons and positrons, fission fragments, neutrons, and y radiation), which typically have very short ranges, the antineutrinos produced in β^{-} decay interact so weakly with ordinary matter that the vast majority pass right out into space. Thus they can be detected in remote terrestrial locations, leading to the possibility of using antineutrino measurements as a probe into radioactive decay occurring within the earth.

In this way further information on the earth's current level of radiogenic heat production has been gleaned from measurements of geoneutrinos in the underground detectors KamLAND (Kamioka Liquid-Scintillator Antineutrino Detector, located near Toyama in Japan) and Borexino (at Laboratori Nazionali del Gran Sasso near L'Aquila, Italy). Using combined electron antineutrino measurements from both KamLAND and Borexino, KamLAND scientists estimated that of a total heat flow from the earth to space of 44.2±1.0TW (cited from Pollack et al. 1993), the decay of ²³⁸U and ²³²Th together contribute 20.0(+8.8,-8.6) TW. They also assert that the decay of ⁴⁰K contributes a total of 4TW (Gando et al. 2011). The Borexino Collaboration (Agostini et al. 2015) concludes from over 2000 days of antineutrino measurements that the radiogenic heat production for U and Th from the present best-fit result is restricted to the range 23-36TW, taking into account the uncertainty on the distribution of heat producing elements inside the earth.

In deducing this result Agostini et al. (2015) have assumed a chondritic mass ratio of uranium and thorium, viz. m(Th)/m(U)=3.9. Furthermore, assuming a chondritic ratio of potassium to uranium, viz $m(K)/m(U)=10^4$, they estimate a total terrestrial radiogenic power P(U+Th+K) of 33(+28,-20) TW. The central value of this estimate is considerably larger than KamLAND's earlier one, but the large margins of uncertainty in both cases mean that

they are not necessarily in conflict. Although these antineutrino-based estimates do not provide decisive answers to questions of present-day radiogenic heating, future developments in detection and measurement technology may give much more accurate estimates than is currently possible.

Conclusions

- 1. The most widely used ocean temperature indicator is the δ^{18} O level, which is recorded in the shells of fossil marine animals throughout the Phanerozoic and in inorganic Precambrian deposits. The trend. notwithstanding interpretive uncertainties and occasional large variations, is of $\delta^{18}O$ increasing through time. Whether this represents falling temperatures, an increasing seawater $\delta^{18}O$ level or diagenetic alteration, the Cenozoic portion of the record appears to be a broadly reliable indicator of temperature and global ice volume. Thus the highest ocean floor temperatures since the Mesozoic were about 12–13°C in the early Eocene. All models of the Flood and its aftermath need to take this limit into significantly higher bulk account: ocean temperatures during the late stages of the Flood or afterwards would conflict with the data.
- 2. Earth's surface heat flow totaling 47 ± 2 TW consists mainly of radiogenic heat from the continental crust and lower mantle, heat flow from the core, and cooling of oceanic lithosphere. Although heatproducing radioactive nuclide inventories can only be assessed indirectly, estimates based on rock compositions and geoneutrino measurements are in broad agreement. Thus the enormous heat load due to AND during the Flood can be estimated with a fair degree of confidence. A rough preliminary estimate based on 600 million years' worth of nuclear decay during the Flood year suggests that, for example, the radiogenic heat generated within a sizable granite body (i.e. more than a few meters across) emplaced before the Flood is times enough more than 13to melt it completely. The present-day pattern of continental heat flow, which is problematic in the uniformitarian paradigm, has been neatly explained by Baumgardner (2000) in terms of a limited net pulse of AND during the Flood. However the question of how almost all of the AND heat in the continental crust was removed has not yet been answered. Thus the challenge for Flood models is to identify and describe a sufficiently strong cooling mechanism to remove all but a tiny fraction of the heat generated by AND, mainly but not exclusively in the earth's continental crust.
- 3. Uniformitarian models of ocean floor cooling are

not entirely satisfactory. However, since ocean floors are geologically young and almost certainly formed during and after the Flood, Flood models face the challenge of explaining how the heat released during their formation was removed without boiling off the oceans. A recent attempt to solve this problem by invoking a transient subsurface heat sink together with rapid oceanic plate movement was unsuccessful, suggesting that neglected but important geophysical effects (e.g. deep hydrothermal convection, heat removal by superheated steam jets, etc.) may need to be included in future.

4. Heterogeneities detected seismically and in other ways exist on a wide range of scales throughout the crust and mantle. Many of these provide evidence of sinking or upwelling motions in the mantle, and some in the deep mantle signify subducted slabs which are unlikely to have retained their identity over many millions of years. Since mantle heterogeneities serve as an archive of the earth's tectonic history, which is closely interwoven with its thermal history, they should be taken into account at some stage in the development of Flood models which seek to address heat problems.

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Nomenclature

- **Benthic**: Benthic organisms (collectively benthos) live at the bottom of oceans or lakes. Organisms living on the bottom, within the surface layer or just below are included in the definition.
- **Conodonts**: Conodonts, or conodont elements, are the "teeth" of eel-like animals, ranging stratigraphically from Cambrian to late Triassic (Sweet and Donoghue 2001); conodont elements are made of calcium phosphate (Donoghue, Forey, and Aldridge 2000).
- **Foraminifera**: Colloquially called forams, foraminifera belong to a class of mostly marine single-celled organisms (technically protists) characterized by streaming granular ectoplasm (for feeding etc.) and which produce an external shell (or "test") of various forms. Foraminiferal tests are most commonly made of calcium carbonate, usually calcite. For further information see Pearson (2012).
- **Photic zone**: The surface layer of the ocean, which is illuminated by sunlight. The uppermost 80 m or so is the euphotic zone, where the light level is sufficient to permit photosynthesis by phytoplankton and plants.
- **Planktonic (or planktic)**: Planktonic/planktic organisms live in open water, generally in the near-surface region. Since they are freely-floating, in the oceans they are carried along by the prevailing near-surface currents. Most are microscopic in size, though larger organisms (e.g. jellyfish) are included in the definition.